

# Siril user's guide

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SIRIL is an abbreviation for nothing except IRIS reversed with a trailing L added for linux(though Siril has nothing linux-specific). This document refers to siril v0.4.

## 1 Who will need Siril?

The amateur astronomer who wants a reasonably user- friendly interface but who also wants to keep an eye on the processings, and keep those processings under control. Basically, it is more IRIS-like than Registax- like.

## 2 Philosophy of operation

You have a working image space, on which basic operations are done in memory. For example, commands like the `soper`, `ioper`, `crop`, `wavelet` commands act on that working space and not on image files.

From the programmer's point of view, this is represented by the `fits` structure pointed to by `com.g`. `com` is an all-purpose structure, instantiated only once and that is basically intended to provide easy access to a set of global variables that may be needed by the different modules of the software; and the working image space is one of those variables.

## 3 Command summary

### 3.1 load

`load filename [rvbag]`

`load` loads the fits image `filename` ; it first attempts to load `filename`, then `filename.fit` and finally `filename.fits`, aborting if none of these are found. This scheme is applicable to every `siril` command implying reading files. Fits headers MIPS-HI and MIPS-LO are read and their values given to the current viewing levels (the second, optional argument is reserved for a future use).

### 3.2 save

`save filename [rvbag]`

`save` writes the currently displayed image to `filename.fit`. Fits headers MIPS-HI and MIPS-LO are added with values corresponding to the current viewing levels (the second, optional argument is reserved for a future use).

### 3.3 ioper commands : `iadd, isub, imul, idiv`

`ioper filename`

`ioper` (i for 'image') commands respectively adds, subtract, multiplies or divides current image by the fits image filename, the resulting image is displayed but not saved.

### 3.4 soper commands : `sadd, ssub, smul, sdiv`

`soper scalar`

`soper` commands ('s' for scalar) respectively adds, subtract, multiplies or divides current image by `scalar`, the resulting image is displayed but not saved.

### 3.5 ioper2 commands : iadd2,isub2,imul2,idiv2

ioper2 `gname outname filename number`

ioper2 commands respectively adds, subtract, multiplies or divides the sequence of `number` images with generic name `gname` by the fits image filename, the resulting images are stored in the corresponding sequence with generic name `outname`.

### 3.6 soper commands : sadd2,ssub2,smul2,sdiv2

soper2 `gname outname scalar number`

soper2 commands respectively adds, subtract, multiplies or divides the sequence of `number` images with generic name `gname` by the given scalar, the resulting images are stored in the corresponding sequence with generic name `outname`.

### 3.7 fdiv

fdiv `filename scalar`

fdiv divides current image by the image `filename` and multiplies the result by `scalar` ; this cant be done with idiv and smul because floats must be used here.

### 3.8 log

log

log computes and applies a logarithmic scale to the current image.

### 3.9 threshlo, threshhi, thresh

These are threshold functions:

- `threshlo 0` replaces negative values with 0s;
- `threshhi 1000` replaces values above 1000 with 1000;
- `thresh 0 1000` does both.

### 3.10 nozero

`nozero level`

`nozero level` replaces null values by `level` values. Useful before an `idiv` or `fdiv` operation.

### 3.11 ddp

`ddp level coef sigma`

`ddp` performs a DDP (digital development processing) as described first by Kunihiro Okano. This implementation is the one described in IRIS:

<http://www.astrosurf.org/buil/us/iris/iris5.htm>

It combines a linear distribution on low levels (below `level`) and a non-linear on high levels. It uses a gaussian filter of sigma `sigma` and multiplies the resulting image by `coef`.

### 3.12 register

`register gename number [x y width height]`

`register` register the given image sequence as a function of the specified image region (either working space mouse selection or as command line parameters, as specified in the prototype. `register` creates a file with calculated image shifts and image quality. This file is named `gename.shift`, and can be user for further operations, typically `composit`, that will read the shift file on the fly to produce an aligned stacking of the initial image set.

### 3.13 `composit` also known as `stack`

```
composit gname number [shiftfile]
```

This command stacks the given image sequence of `n` images to the current working image. If `shiftfile` is specified, `x` and `y` shift values are read on the fly (the sequence must have been registered first); if no `shiftfile` is given, stacking is done with no shift.

### 3.14 `shift`

```
shift x y
```

This command shifts current image by `x` and `y`.

### 3.15 `shift2`

```
shift2 gname outname number x y
```

This command shifts `gname` image sequence by `x` and `y`, outputting the result to the `outname` sequence.

### 3.16 `rshift2`

```
rshift2 gname outname number [shiftfile]
```

This command shifts `gname` image sequence by `x` and `y`, outputting the result to the `outname` sequence. `x` and `y` shift are read on the fly from `shiftfile` (the sequence must have been registered first).

### 3.17 `medstack`

```
medstack gname number
```

This command puts in working image the median stack of the given sequence. `number` is preferably odd.

### 3.18 unsharp, unsharp2

`unsharp sigma coef`

`unsharp` applies to the working image an unsharp mask with sigma `sigma` and coefficient `coef`.

`unsharp2 sigma coef genname outname number`

`unsharp2` is the sequence version of `tt unsharp`, applying `unsharp` to the sequence `genname` the result goes in the sequence `outname`.

### 3.19 gauss

`gauss sigma`

`gauss` performs a gaussian filter with the given sigma.

### 3.20 wavelet

`wavelet plan_number type (1=linear 2=spline)`

computes the wavelet transform on `plan_number` planes using linear (`type = 1`) or bspline (`type=2`) version of the 'a trous' algorithm. The result is stored in a file as a structure containing the planes, ready for weighted reconstruction with `wrecons`.

### 3.21 wrecons

`wrecons c1 c2 ... cn`

`wrecons` reconstruct to current image from the planes previously computed with `wavelet` and weighted with coefficients `c1`, `c2`, ... `cn` according to the number of planes used for wavelet transform.

### 3.22 `animate`

```
animate gname start_image end_image delay stride
```

`animate` animates the `gname` sequence of images, starting at image number `start_image`, ending at `end_image` with an optional delay between images and optional stride (displays 1 image each `stride` images). `animate` loops until interrupted by `ESC` key.